

Carroll Fermions and Supersymmetry

Mojtaba Najafizadeh

Department of Physics, Faculty of Science, Ferdowsi University of Mashhad (FUM)
Mashhad, Iran

School of physics, Institute for Research in Fundamental Sciences (IPM)
Tehran, Iran

School of Particles and Accelerators, Institute for Research in Fundamental Sciences (IPM)

Wednesday Weekly Seminars

Jan 24, 2024



Based on:

K. Koutrolikos & M. Najafizadeh,

“Super-Carrollian and super-Galilean field theories”,

[Phys. Rev. D **108**, 125014 \(2023\)](#), [arXiv:2309.16786 \[hep-th\]](#)

Terminology

Terminology:

Ann. Inst. Henri Poincaré,
Vol. III, n° 1, 1965, p. 1-12.

Section A :
Physique théorique.

**Une nouvelle limite non-relativiste
du groupe de Poincaré**

par

Jean-Marc LÉVY-LEBLOND
(Laboratoire de Physique Théorique, Orsay).

ABSTRACT. — It is shown that, for the Galilean approximation to the Poincaré group to be valid, not only do we have to consider pure Lorentz transformation with low velocities, but also great time-like intervals. Under the same conditions, the transformation properties of great space-like inter-

Terminology:

6

JEAN-MARC LEVY-LEBLOND

P_0 le générateur des translations de temps. Il est instructif de comparer les algèbres de Lie du groupe de Carroll et des groupes de Poincaré et Galilée. On a respectivement :

$$[J_i, J_j] = i\varepsilon_{ijk}J_k$$

$$[J_i, K_j] = i\varepsilon_{ijk}K_k$$

$$[K_i, K_j] = 0$$

$$[J_i, P_j] = i\varepsilon_{ijk}P_k$$

$$[K_i, P_j] = 0$$

$$[J_i, P_0] = 0$$

$$[K_i, P_0] = iP_i$$

$$[P_i, P_j] = 0$$

$$[P_i, P_0] = 0$$

Galilée

$$[J_i, J_j] = i\varepsilon_{ijk}J_k$$

$$[J_i, K_j] = i\varepsilon_{ijk}K_k$$

$$[K_i, K_j] = -i\varepsilon_{ijk}J_k$$

$$[J_i, P_j] = i\varepsilon_{ijk}P_k$$

$$[K_i, P_j] = i\delta_{ij}P_0$$

$$[J_i, P_0] = 0$$

$$[K_i, P_0] = iP_i$$

$$[P_i, P_j] = 0$$

$$[P_i, P_0] = 0$$

Poincaré

$$[J_i, J_j] = i\varepsilon_{ijk}J_k$$

$$[J_i, K_j] = i\varepsilon_{ijk}K_k$$

$$[K_i, K_j] = 0$$

$$[J_i, P_j] = i\varepsilon_{ijk}P_k$$

$$[K_i, P_j] = i\delta_{ij}P_0$$

$$[J_i, P_0] = 0$$

$$[K_i, P_0] = 0$$

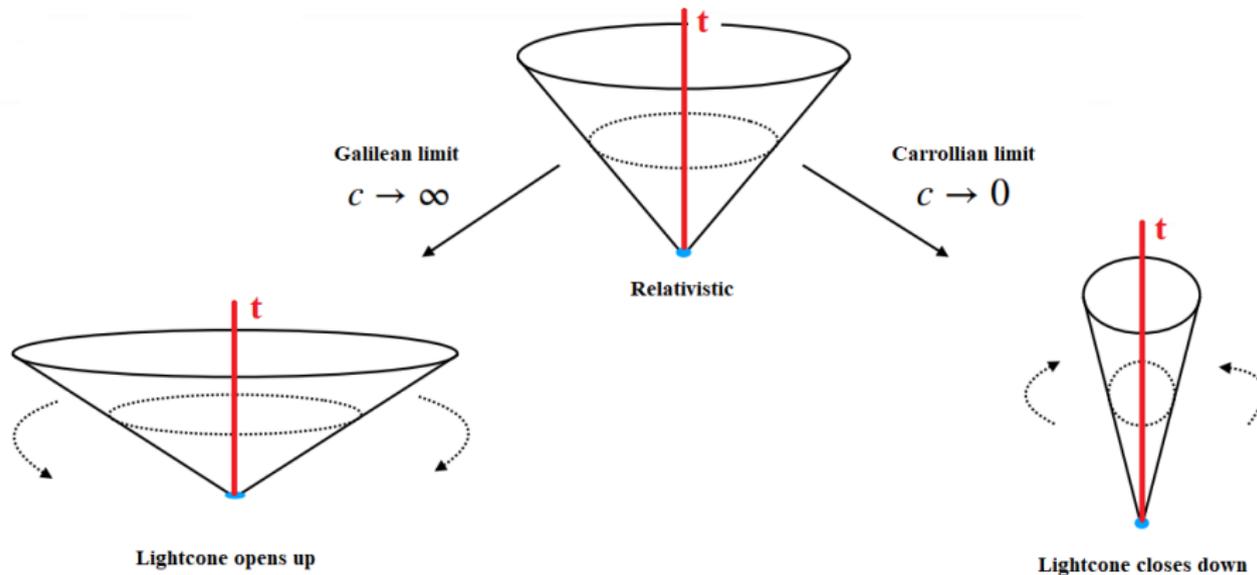
$$[P_i, P_j] = 0$$

$$[P_i, P_0] = 0$$

Carroll

On voit clairement sur ces expressions que substituant :

Terminology:



Galilei limit: $\frac{v}{c} \rightarrow 0 \iff c \rightarrow \infty$

Carroll limit: $\frac{v}{c} \rightarrow \infty \iff c \rightarrow 0$

Outline:

Motivation

Carroll Particles

Carroll Algebra

Carroll Scalar Field Theories ($s = 0$)

Carroll Fermion Field Theories ($s = \frac{1}{2}$) ✓

On Supersymmetry

Carrollian Supersymmetry ($0, \frac{1}{2}$) ✓

Conclusion

Motivation

Motivation

Flat space holography:

Conformal Carroll field theory might be dual to quantum gravity in asymptotically flat spacetime [S. Pasterski (2021), L. Donnay (2023), ...]

Motivation

Flat space holography:

Conformal Carroll field theory might be dual to quantum gravity in asymptotically flat spacetime [S. Pasterski (2021), L. Donnay (2023), ...]

Cosmology and dark energy:

Carroll symmetry might be relevant for de Sitter cosmology and inflation [J. de Boer, J. Hartong (2022), ...]

Flat space holography:

Conformal Carroll field theory might be dual to quantum gravity in asymptotically flat spacetime [S. Pasterski (2021), L. Donnay (2023), ...]

Cosmology and dark energy:

Carroll symmetry might be relevant for de Sitter cosmology and inflation [J. de Boer, J. Hartong (2022), ...]

Carroll gravity:

M. Henneaux (1979), N. A. Obers (2022), D. Grumiller (2023), ...

Flat space holography:

Conformal Carroll field theory might be dual to quantum gravity in asymptotically flat spacetime [S. Pasterski (2021), L. Donnay (2023), ...]

Cosmology and dark energy:

Carroll symmetry might be relevant for de Sitter cosmology and inflation [J. de Boer, J. Hartong (2022), ...]

Carroll gravity:

M. Henneaux (1979), N. A. Obers (2022), D. Grumiller (2023), ...

String theory:

Carroll symmetries arise in the tensionless limit of string theory [A. Bagchi (2016), ...]

Flat space holography:

Conformal Carroll field theory might be dual to quantum gravity in asymptotically flat spacetime [S. Pasterski (2021), L. Donnay (2023), ...]

Cosmology and dark energy:

Carroll symmetry might be relevant for de Sitter cosmology and inflation [J. de Boer, J. Hartong (2022), ...]

Carroll gravity:

M. Henneaux (1979), N. A. Obers (2022), D. Grumiller (2023), ...

String theory:

Carroll symmetries arise in the tensionless limit of string theory [A. Bagchi (2016), ...]

Carroll symmetry in hydrodynamics:

L. Ciambelli, C. Marteau (2018), ...

Carroll Algebra

Carroll Algebra

Recall the Poincaré algebra

$$[P_\mu, P_\nu] = 0$$

$$[J_{\mu\nu}, P_\rho] = \eta_{\mu\rho} P_\nu - \eta_{\nu\rho} P_\mu$$

$$[J_{\mu\nu}, J_{\rho\sigma}] = \eta_{\mu\rho} J_{\nu\sigma} - \eta_{\mu\sigma} J_{\nu\rho} - \eta_{\nu\rho} J_{\mu\sigma} + \eta_{\nu\sigma} J_{\mu\rho}$$

Carroll Algebra

Recall the Poincaré algebra

$$[P_\mu, P_\nu] = 0$$

$$[J_{\mu\nu}, P_\rho] = \eta_{\mu\rho} P_\nu - \eta_{\nu\rho} P_\mu$$

$$[J_{\mu\nu}, J_{\rho\sigma}] = \eta_{\mu\rho} J_{\nu\sigma} - \eta_{\mu\sigma} J_{\nu\rho} - \eta_{\nu\rho} J_{\mu\sigma} + \eta_{\nu\sigma} J_{\mu\rho}$$

in which

$$P_\mu := \partial_\mu \quad \text{Translation generators} \quad \left\{ \begin{array}{ll} P_0 & \text{Hamiltonian} \\ P_i & \text{Momentum} \end{array} \right.$$

$$J_{\mu\nu} := x_\mu \partial_\nu - x_\nu \partial_\mu \quad \text{Lorentz generators} \quad \left\{ \begin{array}{ll} J_{i0} & \text{Boosts} \\ J_{ij} & \text{Rotations} \end{array} \right.$$

are ten Poincaré generators in 4d spacetime.

Carroll Algebra

Poincare generators, P_0 , P_i , J_{i0} , J_{ij} , including c (i.e. $x^\mu = (ct, x^i)$) become:

$$P_0 = \frac{1}{c} \partial_t \quad P_i = \partial_i \quad J_{i0} = x_i \frac{1}{c} \partial_t + ct \partial_i \quad J_{ij} = x_i \partial_j - x_j \partial_i$$

Carroll Algebra

Poincare generators, P_0 , P_i , J_{i0} , J_{ij} , including c (i.e. $x^\mu = (ct, x^i)$) become:

$$P_0 = \frac{1}{c} \partial_t \quad P_i = \partial_i \quad J_{i0} = x_i \frac{1}{c} \partial_t + ct \partial_i \quad J_{ij} = x_i \partial_j - x_j \partial_i$$

Rescaling the generators, one gets

$$c P_0 \rightarrow H = \partial_t \quad c J_{i0} \rightarrow B_i = x_i \partial_t + c^2 t \partial_i$$

Carroll Algebra

Poincare generators, P_0 , P_i , J_{i0} , J_{ij} , including c (i.e. $x^\mu = (ct, x^i)$) become:

$$P_0 = \frac{1}{c} \partial_t \quad P_i = \partial_i \quad J_{i0} = x_i \frac{1}{c} \partial_t + ct \partial_i \quad J_{ij} = x_i \partial_j - x_j \partial_i$$

Rescaling the generators, one gets

$$c P_0 \rightarrow H = \partial_t \quad c J_{i0} \rightarrow B_i = x_i \partial_t + c^2 t \partial_i$$

Taking the Carroll limit $c \rightarrow 0$, one obtains the **Carroll generators**:

$$H = \partial_t \quad P_i = \partial_i \quad B_i = x_i \partial_t \quad J_{ij} = x_i \partial_j - x_j \partial_i$$

Carroll Algebra

Poincare generators, P_0 , P_i , J_{i0} , J_{ij} , including c (i.e. $x^\mu = (ct, x^i)$) become:

$$P_0 = \frac{1}{c} \partial_t \quad P_i = \partial_i \quad J_{i0} = x_i \frac{1}{c} \partial_t + ct \partial_i \quad J_{ij} = x_i \partial_j - x_j \partial_i$$

Rescaling the generators, one gets

$$c P_0 \rightarrow H = \partial_t \quad c J_{i0} \rightarrow B_i = x_i \partial_t + c^2 t \partial_i$$

Taking the Carroll limit $c \rightarrow 0$, one obtains the **Carroll generators**:

$$H = \partial_t \quad P_i = \partial_i \quad B_i = x_i \partial_t \quad J_{ij} = x_i \partial_j - x_j \partial_i$$

satisfying the **Carroll algebra**:

$$\begin{aligned} [P_i, B_j] &= \delta_{ij} H \\ [J_{ij}, P_k] &= \delta_{ik} P_j - \delta_{jk} P_i \\ [J_{ij}, B_k] &= \delta_{ik} B_j - \delta_{jk} B_i \\ [J_{ij}, J_{kl}] &= \delta_{ik} J_{jl} - \delta_{jk} J_{il} + \delta_{jl} J_{ik} - \delta_{il} J_{jk} \end{aligned}$$

Carroll Particles

Carroll Particles

Consider a free relativistic particle with energy E , mass m and momentum \vec{p} :

$$\vec{p} = \gamma m \vec{v} \qquad E = \gamma m c^2 \qquad E^2 = \vec{p}^2 c^2 + m^2 c^4$$

Carroll Particles

Consider a free relativistic particle with energy E , mass m and momentum \vec{p} :

$$\vec{p} = \gamma m \vec{v} \qquad E = \gamma m c^2 \qquad E^2 = \vec{p}^2 c^2 + m^2 c^4$$

where

$$\gamma = \frac{1}{\sqrt{1 - \frac{v^2}{c^2}}}$$

Carroll Particles

Consider a free relativistic particle with energy E , mass m and momentum \vec{p} :

$$\vec{p} = \gamma m \vec{v} \qquad E = \gamma m c^2 \qquad E^2 = \vec{p}^2 c^2 + m^2 c^4$$

where

$$\gamma = \frac{1}{\sqrt{1 - \frac{v^2}{c^2}}} \xrightarrow{c \rightarrow 0} ?$$

is not well-defined in the Carroll limit $c \rightarrow 0$.

Carroll Particles

Consider a free relativistic particle with energy E , mass m and momentum \vec{p} :

$$\vec{p} = \gamma m \vec{v} \qquad E = \gamma m c^2 \qquad E^2 = \vec{p}^2 c^2 + m^2 c^4$$

where

$$\gamma = \frac{1}{\sqrt{1 - \frac{v^2}{c^2}}} \xrightarrow{c \rightarrow 0} ?$$

is not well-defined in the Carroll limit $c \rightarrow 0$. There are two options:

$$\left\{ \begin{array}{l} \vec{v} = 0 \quad \rightarrow \quad \gamma = 1 \quad \rightarrow \quad \vec{p} = 0 \quad E = mc^2 \quad \text{electric Carroll particles} \end{array} \right.$$

Carroll Particles

Consider a free relativistic particle with energy E , mass m and momentum \vec{p} :

$$\vec{p} = \gamma m \vec{v} \quad E = \gamma m c^2 \quad E^2 = \vec{p}^2 c^2 + m^2 c^4$$

where

$$\gamma = \frac{1}{\sqrt{1 - \frac{v^2}{c^2}}} \xrightarrow{c \rightarrow 0} ?$$

is not well-defined in the Carroll limit $c \rightarrow 0$. There are two options:

$$\left\{ \begin{array}{l} \vec{v} = 0 \quad \rightarrow \quad \gamma = 1 \quad \rightarrow \quad \vec{p} = 0 \quad E = mc^2 \quad \text{electric Carroll particles} \\ v^2 > c^2 \quad \rightarrow \quad \gamma = ic/v \quad \rightarrow \quad \vec{p} = mc \hat{v} \quad E = 0 \quad \text{magnetic Carroll particles} \end{array} \right.$$

Carroll Particles

Consider a free relativistic particle with energy E , mass m and momentum \vec{p} :

$$\vec{p} = \gamma m \vec{v} \qquad E = \gamma m c^2 \qquad E^2 = \vec{p}^2 c^2 + m^2 c^4$$

where

$$\gamma = \frac{1}{\sqrt{1 - \frac{v^2}{c^2}}} \xrightarrow{c \rightarrow 0} ?$$

is not well-defined in the Carroll limit $c \rightarrow 0$. There are two options:

$$\left\{ \begin{array}{llllll} \vec{v} = 0 & \rightarrow & \gamma = 1 & \rightarrow & \vec{p} = 0 & E = mc^2 & \text{electric Carroll particles} \\ v^2 > c^2 & \rightarrow & \gamma = ic/v & \rightarrow & \vec{p} = mc \hat{v} & E = 0 & \text{magnetic Carroll particles} \end{array} \right.$$

Jan de Boer et al. [arxiv:2110.02319]

Carroll Scalar Field Theories ($s = 0$)

There are 4 methods in the literature:

- Lagrange multiplier method ✓
- Hamiltonian method ✓
- Field expansion method ✓
- Seed Lagrangian method

Carroll Scalar Field Theories

Let us consider the Lagrangian of a relativistic massless scalar field

$$\mathcal{L} = -\frac{1}{2} \partial_\mu \phi \partial^\mu \phi$$

Carroll Scalar Field Theories

Let us consider the Lagrangian of a relativistic massless scalar field

$$\mathcal{L} = -\frac{1}{2} \partial_\mu \phi \partial^\mu \phi$$

including the speed of light c

$$\mathcal{L} = \frac{1}{2c^2} (\partial_t \phi)^2 - \frac{1}{2} (\partial_i \phi)^2$$

Carroll Scalar Field Theories

Let us consider the Lagrangian of a relativistic massless scalar field

$$\mathcal{L} = -\frac{1}{2} \partial_\mu \phi \partial^\mu \phi$$

including the speed of light c

$$\mathcal{L} = \frac{1}{2c^2} (\partial_t \phi)^2 - \frac{1}{2} (\partial_i \phi)^2$$

by rescaling the field $\phi \rightarrow c\phi$, one has

$$\mathcal{L} = \frac{1}{2} (\partial_t \phi)^2 - \frac{1}{2} c^2 (\partial_i \phi)^2$$

Carroll Scalar Field Theories

Let us consider the Lagrangian of a relativistic massless scalar field

$$\mathcal{L} = -\frac{1}{2} \partial_\mu \phi \partial^\mu \phi$$

including the speed of light c

$$\mathcal{L} = \frac{1}{2c^2} (\partial_t \phi)^2 - \frac{1}{2} (\partial_i \phi)^2$$

by rescaling the field $\phi \rightarrow c\phi$, one has

$$\mathcal{L} = \frac{1}{2} (\partial_t \phi)^2 - \frac{1}{2} c^2 (\partial_i \phi)^2$$

Taking the Carroll limit $c \rightarrow 0$,

Carroll Scalar Field Theories

Let us consider the Lagrangian of a relativistic massless scalar field

$$\mathcal{L} = -\frac{1}{2} \partial_\mu \phi \partial^\mu \phi$$

including the speed of light c

$$\mathcal{L} = \frac{1}{2c^2} (\partial_t \phi)^2 - \frac{1}{2} (\partial_i \phi)^2$$

by rescaling the field $\phi \rightarrow c\phi$, one has

$$\mathcal{L} = \frac{1}{2} (\partial_t \phi)^2 - \frac{1}{2} c^2 (\partial_i \phi)^2$$

Taking the Carroll limit $c \rightarrow 0$, one obtains

Electric Carroll scalar field:

$$\mathcal{L}_e = \frac{1}{2} (\partial_t \phi)^2$$

Carroll Scalar Field Theories

Let us consider the Lagrangian of a relativistic massless scalar field

$$\mathcal{L} = -\frac{1}{2} \partial_\mu \phi \partial^\mu \phi$$

including the speed of light c

$$\mathcal{L} = \frac{1}{2c^2} (\partial_t \phi)^2 - \frac{1}{2} (\partial_i \phi)^2$$

by rescaling the field $\phi \rightarrow c\phi$, one has

$$\mathcal{L} = \frac{1}{2} (\partial_t \phi)^2 - \frac{1}{2} c^2 (\partial_i \phi)^2$$

Taking the Carroll limit $c \rightarrow 0$, one obtains

Electric Carroll scalar field:

$$\mathcal{L}_e = \frac{1}{2} (\partial_t \phi)^2$$

- mass term can be added

Carroll Scalar Field Theories

Let us consider the Lagrangian of a relativistic massless scalar field

$$\mathcal{L} = -\frac{1}{2} \partial_\mu \phi \partial^\mu \phi$$

including the speed of light c

$$\mathcal{L} = \frac{1}{2c^2} (\partial_t \phi)^2 - \frac{1}{2} (\partial_i \phi)^2$$

by rescaling the field $\phi \rightarrow c\phi$, one has

$$\mathcal{L} = \frac{1}{2} (\partial_t \phi)^2 - \frac{1}{2} c^2 (\partial_i \phi)^2$$

Taking the Carroll limit $c \rightarrow 0$, one obtains

Electric Carroll scalar field:

$$\mathcal{L}_e = \frac{1}{2} (\partial_t \phi)^2$$

- mass term can be added

Carroll Scalar Field Theories

Let us consider again the Lagrangian of a relativistic massless scalar field

$$\mathcal{L} = -\frac{1}{2} \partial_\mu \phi \partial^\mu \phi$$

Carroll Scalar Field Theories

Let us consider again the Lagrangian of a relativistic massless scalar field

$$\mathcal{L} = -\frac{1}{2} \partial_\mu \phi \partial^\mu \phi$$

including the speed of light c

$$\mathcal{L} = \frac{1}{2c^2} (\partial_t \phi)^2 - \frac{1}{2} (\partial_i \phi)^2 \quad (1)$$

Carroll Scalar Field Theories

Let us consider again the Lagrangian of a relativistic massless scalar field

$$\mathcal{L} = -\frac{1}{2} \partial_\mu \phi \partial^\mu \phi$$

including the speed of light c

$$\mathcal{L} = \frac{1}{2c^2} (\partial_t \phi)^2 - \frac{1}{2} (\partial_i \phi)^2 \quad (1)$$

This Lagrangian has also an alternative by adding a Lagrange multiplier χ :

$$\mathcal{L}' = -\frac{1}{2} c^2 \chi^2 + \chi \partial_t \phi - \frac{1}{2} (\partial_i \phi)^2 \quad -c^2 \chi + \partial_t \phi = 0 \quad (2)$$

Carroll Scalar Field Theories

Let us consider again the Lagrangian of a relativistic massless scalar field

$$\mathcal{L} = -\frac{1}{2} \partial_\mu \phi \partial^\mu \phi$$

including the speed of light c

$$\mathcal{L} = \frac{1}{2c^2} (\partial_t \phi)^2 - \frac{1}{2} (\partial_i \phi)^2 \quad (1)$$

This Lagrangian has also an alternative by adding a Lagrange multiplier χ :

$$\mathcal{L}' = -\frac{1}{2} c^2 \chi^2 + \chi \partial_t \phi - \frac{1}{2} (\partial_i \phi)^2 \quad -c^2 \chi + \partial_t \phi = 0 \quad (2)$$

Taking the Carroll limit $c \rightarrow 0$,

Carroll Scalar Field Theories

Let us consider again the Lagrangian of a relativistic massless scalar field

$$\mathcal{L} = -\frac{1}{2} \partial_\mu \phi \partial^\mu \phi$$

including the speed of light c

$$\mathcal{L} = \frac{1}{2c^2} (\partial_t \phi)^2 - \frac{1}{2} (\partial_i \phi)^2 \quad (1)$$

This Lagrangian has also an alternative by adding a Lagrange multiplier χ :

$$\mathcal{L}' = -\frac{1}{2} c^2 \chi^2 + \chi \partial_t \phi - \frac{1}{2} (\partial_i \phi)^2 \quad -c^2 \chi + \partial_t \phi = 0 \quad (2)$$

Taking the Carroll limit $c \rightarrow 0$, one obtains

Magnetic Carroll scalar field:

$$\mathcal{L}_m = \chi \partial_t \phi - \frac{1}{2} (\partial_i \phi)^2$$

Carroll Scalar Field Theories

Let us consider again the Lagrangian of a relativistic massless scalar field

$$\mathcal{L} = -\frac{1}{2} \partial_\mu \phi \partial^\mu \phi$$

including the speed of light c

$$\mathcal{L} = \frac{1}{2c^2} (\partial_t \phi)^2 - \frac{1}{2} (\partial_i \phi)^2 \quad (1)$$

This Lagrangian has also an alternative by adding a Lagrange multiplier χ :

$$\mathcal{L}' = -\frac{1}{2} c^2 \chi^2 + \chi \partial_t \phi - \frac{1}{2} (\partial_i \phi)^2 \quad -c^2 \chi + \partial_t \phi = 0 \quad (2)$$

Taking the Carroll limit $c \rightarrow 0$, one obtains

Magnetic Carroll scalar field:

$$\mathcal{L}_m = \chi \partial_t \phi - \frac{1}{2} (\partial_i \phi)^2$$

- mass term can be added as a tachyon

Carroll Scalar Field Theories

Let us consider again the Lagrangian of a relativistic massless scalar field

$$\mathcal{L} = -\frac{1}{2} \partial_\mu \phi \partial^\mu \phi$$

including the speed of light c

$$\mathcal{L} = \frac{1}{2c^2} (\partial_t \phi)^2 - \frac{1}{2} (\partial_i \phi)^2 \quad (1)$$

This Lagrangian has also an alternative by adding a Lagrange multiplier χ :

$$\mathcal{L}' = -\frac{1}{2} c^2 \chi^2 + \chi \partial_t \phi - \frac{1}{2} (\partial_i \phi)^2 \quad -c^2 \chi + \partial_t \phi = 0 \quad (2)$$

Taking the Carroll limit $c \rightarrow 0$, one obtains

Magnetic Carroll scalar field:

$$\mathcal{L}_m = \chi \partial_t \phi - \frac{1}{2} (\partial_i \phi)^2$$

- mass term can be added as a tachyon

J. de Boer (2022), E. A. Bergshoeff (2022)

Carroll Scalar Field Theories

Hamiltonian method: Let us consider a massless scalar field Lagrangian in which the canonical momentum conjugate to the field reads

$$\pi_\phi = \frac{\partial \mathcal{L}}{\partial \dot{\phi}} = \frac{1}{c^2} \dot{\phi}, \quad \dot{\phi} \equiv \partial_t \phi$$

Carroll Scalar Field Theories

Hamiltonian method: Let us consider a massless scalar field Lagrangian in which the canonical momentum conjugate to the field reads

$$\pi_\phi = \frac{\partial \mathcal{L}}{\partial \dot{\phi}} = \frac{1}{c^2} \dot{\phi}, \quad \dot{\phi} \equiv \partial_t \phi$$

and then the Hamiltonian density becomes

$$\mathcal{H} = \pi_\phi \dot{\phi} - \mathcal{L} = \frac{1}{2} \left[c^2 (\pi_\phi)^2 + (\partial_i \phi)^2 \right]$$

Carroll Scalar Field Theories

Hamiltonian method: Let us consider a massless scalar field Lagrangian in which the canonical momentum conjugate to the field reads

$$\pi_\phi = \frac{\partial \mathcal{L}}{\partial \dot{\phi}} = \frac{1}{c^2} \dot{\phi}, \quad \dot{\phi} \equiv \partial_t \phi$$

and then the Hamiltonian density becomes

$$\mathcal{H} = \pi_\phi \dot{\phi} - \mathcal{L} = \frac{1}{2} \left[c^2 (\pi_\phi)^2 + (\partial_i \phi)^2 \right]$$

Accordingly, the Hamiltonian action becomes

$$S[\phi, \pi_\phi] = \int dt d^3x \left\{ \pi_\phi \dot{\phi} - \mathcal{H} \right\} = \int dt d^3x \left\{ \pi_\phi \dot{\phi} - \frac{1}{2} c^2 (\pi_\phi)^2 - \frac{1}{2} (\partial_i \phi)^2 \right\}$$

Carroll Scalar Field Theories

Hamiltonian method: Let us consider a massless scalar field Lagrangian in which the canonical momentum conjugate to the field reads

$$\pi_\phi = \frac{\partial \mathcal{L}}{\partial \dot{\phi}} = \frac{1}{c^2} \dot{\phi}, \quad \dot{\phi} \equiv \partial_t \phi$$

and then the Hamiltonian density becomes

$$\mathcal{H} = \pi_\phi \dot{\phi} - \mathcal{L} = \frac{1}{2} \left[c^2 (\pi_\phi)^2 + (\partial_i \phi)^2 \right]$$

Accordingly, the Hamiltonian action becomes

$$S[\phi, \pi_\phi] = \int dt d^3x \left\{ \pi_\phi \dot{\phi} - \mathcal{H} \right\} = \int dt d^3x \left\{ \pi_\phi \dot{\phi} - \frac{1}{2} c^2 (\pi_\phi)^2 - \frac{1}{2} (\partial_i \phi)^2 \right\}$$

magnetic: $c \rightarrow 0$

$$S_m = \int dt d^3x \left\{ \pi_\varphi \dot{\phi} - \frac{1}{2} (\partial_i \phi)^2 \right\}$$

Carroll Scalar Field Theories

Hamiltonian method: Let us consider a massless scalar field Lagrangian in which the canonical momentum conjugate to the field reads

$$\pi_\phi = \frac{\partial \mathcal{L}}{\partial \dot{\phi}} = \frac{1}{c^2} \dot{\phi}, \quad \dot{\phi} \equiv \partial_t \phi$$

and then the Hamiltonian density becomes

$$\mathcal{H} = \pi_\phi \dot{\phi} - \mathcal{L} = \frac{1}{2} \left[c^2 (\pi_\phi)^2 + (\partial_i \phi)^2 \right]$$

Accordingly, the Hamiltonian action becomes

$$S[\phi, \pi_\phi] = \int dt d^3x \left\{ \pi_\phi \dot{\phi} - \mathcal{H} \right\} = \int dt d^3x \left\{ \pi_\phi \dot{\phi} - \frac{1}{2} c^2 (\pi_\phi)^2 - \frac{1}{2} (\partial_i \phi)^2 \right\}$$

$$\left\{ \begin{array}{l} \text{magnetic: } c \rightarrow 0 \\ \text{electric: } \phi \rightarrow c\phi, \pi \rightarrow \frac{1}{c}\pi, c \rightarrow 0 \end{array} \right. \quad \left\{ \begin{array}{l} S_m = \int dt d^3x \left\{ \pi_\phi \dot{\phi} - \frac{1}{2} (\partial_i \phi)^2 \right\} \\ S_e = \int dt d^3x \left\{ \pi_\phi \dot{\phi} - \frac{1}{2} (\pi_\phi)^2 \right\} \end{array} \right.$$

Carroll Scalar Field Theories

Hamiltonian method: Let us consider a massless scalar field Lagrangian in which the canonical momentum conjugate to the field reads

$$\pi_\phi = \frac{\partial \mathcal{L}}{\partial \dot{\phi}} = \frac{1}{c^2} \dot{\phi}, \quad \dot{\phi} \equiv \partial_t \phi$$

and then the Hamiltonian density becomes

$$\mathcal{H} = \pi_\phi \dot{\phi} - \mathcal{L} = \frac{1}{2} \left[c^2 (\pi_\phi)^2 + (\partial_i \phi)^2 \right]$$

Accordingly, the Hamiltonian action becomes

$$S[\phi, \pi_\phi] = \int dt d^3x \left\{ \pi_\phi \dot{\phi} - \mathcal{H} \right\} = \int dt d^3x \left\{ \pi_\phi \dot{\phi} - \frac{1}{2} c^2 (\pi_\phi)^2 - \frac{1}{2} (\partial_i \phi)^2 \right\}$$

$$\left\{ \begin{array}{l} \text{magnetic: } c \rightarrow 0 \\ \text{electric: } \phi \rightarrow c\phi, \pi \rightarrow \frac{1}{c}\pi, c \rightarrow 0 \end{array} \right. \quad \left\{ \begin{array}{l} S_m = \int dt d^3x \left\{ \pi_\phi \dot{\phi} - \frac{1}{2} (\partial_i \phi)^2 \right\} \\ S_e = \int dt d^3x \left\{ \pi_\phi \dot{\phi} - \frac{1}{2} (\pi_\phi)^2 \right\} \end{array} \right.$$

M. Henneaux (2021)

Carroll Scalar Field Theories

- Under a Lorentz boost transformation, the scalar field Φ transforms as

$$\delta \Phi = i \beta_i J^{0i} \Phi = \beta_i \left(x^0 \partial^i - x^i \partial^0 \right) \Phi = \left(ct \beta^i \partial_i + \frac{1}{c} \beta^i x_i \partial_t \right) \Phi$$

with β^i being the Lorentz boost parameter, and the Lagrangian transforms into a total derivative.

Carroll Scalar Field Theories

- Under a Lorentz boost transformation, the scalar field Φ transforms as

$$\delta \Phi = i \beta_i J^{0i} \Phi = \beta_i \left(x^0 \partial^i - x^i \partial^0 \right) \Phi = \left(ct \beta^i \partial_i + \frac{1}{c} \beta^i x_i \partial_t \right) \Phi$$

with β^i being the Lorentz boost parameter, and the Lagrangian transforms into a total derivative.

Carroll boost transformations:

- In the Lorentz boost trans., by defining $\beta^i = c b^i$, with b^i being the Carroll boost parameter, and then taking the limit $c \rightarrow 0$, one gets

Carroll Scalar Field Theories

- Under a Lorentz boost transformation, the scalar field Φ transforms as

$$\delta \Phi = i \beta_i J^{0i} \Phi = \beta_i \left(x^0 \partial^i - x^i \partial^0 \right) \Phi = \left(ct \beta^i \partial_i + \frac{1}{c} \beta^i x_i \partial_t \right) \Phi$$

with β^i being the Lorentz boost parameter, and the Lagrangian transforms into a total derivative.

Carroll boost transformations:

- In the Lorentz boost trans., by defining $\beta^i = c b^i$, with b^i being the Carroll boost parameter, and then taking the limit $c \rightarrow 0$, one gets

$$\text{Carroll boost trans.: } \begin{cases} \delta \phi = b^i x_i \partial_t \phi \\ \delta \chi = b^i x_i \partial_t \chi + b^i \partial_i \phi \end{cases}$$

Carroll Fermion Field Theories ($s = \frac{1}{2}$)

Carroll Fermion Field Theories

Let us consider the Lagrangian density of a massless Dirac field including c

$$\mathcal{L} = -\bar{\Psi} (\gamma^\mu \partial_\mu) \Psi = -\bar{\Psi} \left(\frac{1}{c} \gamma^0 \partial_t + \gamma^i \partial_i \right) \Psi$$

Carroll Fermion Field Theories

Let us consider the Lagrangian density of a massless Dirac field including c

$$\mathcal{L} = -\bar{\Psi} (\gamma^\mu \partial_\mu) \Psi = -\bar{\Psi} \left(\frac{1}{c} \gamma^0 \partial_t + \gamma^i \partial_i \right) \Psi$$

We then perform an expansion around $c = 0$ in the Dirac field

$$\Psi = c^\beta \left(\psi + c\eta + \mathcal{O}(c^2) \right)$$

for some β ,

Carroll Fermion Field Theories

Let us consider the Lagrangian density of a massless Dirac field including c

$$\mathcal{L} = -\bar{\Psi} (\gamma^\mu \partial_\mu) \Psi = -\bar{\Psi} \left(\frac{1}{c} \gamma^0 \partial_t + \gamma^i \partial_i \right) \Psi$$

We then perform an expansion around $c = 0$ in the Dirac field

$$\Psi = c^\beta \left(\psi + c \eta + \mathcal{O}(c^2) \right)$$

for some β , and introduce \mathcal{L}_e and \mathcal{L}_m through the Lagrangian density

$$\mathcal{L} = c^{2\beta-1} \left(\mathcal{L}_e + c \mathcal{L}_m + \mathcal{O}(c^2) \right)$$

Carroll Fermion Field Theories

Let us consider the Lagrangian density of a massless Dirac field including c

$$\mathcal{L} = -\bar{\Psi} (\gamma^\mu \partial_\mu) \Psi = -\bar{\Psi} \left(\frac{1}{c} \gamma^0 \partial_t + \gamma^i \partial_i \right) \Psi$$

We then perform an expansion around $c = 0$ in the Dirac field

$$\Psi = c^\beta \left(\psi + c \eta + \mathcal{O}(c^2) \right)$$

for some β , and introduce \mathcal{L}_e and \mathcal{L}_m through the Lagrangian density

$$\mathcal{L} = c^{2\beta-1} \left(\mathcal{L}_e + c \mathcal{L}_m + \mathcal{O}(c^2) \right)$$

By substituting the field expansion into the Lagrangian and comparing with the latter

Carroll Fermion Field Theories

Let us consider the Lagrangian density of a massless Dirac field including c

$$\mathcal{L} = -\bar{\Psi} (\gamma^\mu \partial_\mu) \Psi = -\bar{\Psi} \left(\frac{1}{c} \gamma^0 \partial_t + \gamma^i \partial_i \right) \Psi$$

We then perform an expansion around $c = 0$ in the Dirac field

$$\Psi = c^\beta \left(\psi + c \eta + \mathcal{O}(c^2) \right)$$

for some β , and introduce \mathcal{L}_e and \mathcal{L}_m through the Lagrangian density

$$\mathcal{L} = c^{2\beta-1} \left(\mathcal{L}_e + c \mathcal{L}_m + \mathcal{O}(c^2) \right)$$

By substituting the field expansion into the Lagrangian and comparing with the latter

electric Carroll Dirac: $\mathcal{L}_e = -\bar{\psi} \gamma^0 \partial_t \psi$

magnetic Carroll Dirac: $\mathcal{L}_m = - \left(\bar{\psi} \gamma^i \partial_i \psi + \bar{\eta} \gamma^0 \partial_t \psi + \bar{\psi} \gamma^0 \partial_t \eta \right)$

Carroll Fermion Field Theories

Let us consider the Lagrangian density of a massless Dirac field including c

$$\mathcal{L} = -\bar{\Psi} (\gamma^\mu \partial_\mu) \Psi = -\bar{\Psi} \left(\frac{1}{c} \gamma^0 \partial_t + \gamma^i \partial_i \right) \Psi$$

We then perform an expansion around $c = 0$ in the Dirac field

$$\Psi = c^\beta \left(\psi + c \eta + \mathcal{O}(c^2) \right)$$

for some β , and introduce \mathcal{L}_e and \mathcal{L}_m through the Lagrangian density

$$\mathcal{L} = c^{2\beta-1} \left(\mathcal{L}_e + c \mathcal{L}_m + \mathcal{O}(c^2) \right)$$

By substituting the field expansion into the Lagrangian and comparing with the latter

$$\text{electric Carroll Dirac: } \mathcal{L}_e = -\bar{\psi} \gamma^0 \partial_t \psi$$

$$\text{magnetic Carroll Dirac: } \mathcal{L}_m = - \left(\bar{\psi} \gamma^i \partial_i \psi + \bar{\eta} \gamma^0 \partial_t \psi + \bar{\psi} \gamma^0 \partial_t \eta \right)$$

[K. Koutrolikos, M.N., PRD (2023)]

Carroll Fermion Field Theories

These Lagrangians are invariant under the following Carroll boost transformations:

$$\text{Carroll boost trans.: } \begin{cases} \delta\psi = b^i x_i \partial_t \psi \\ \delta\eta = b^i x_i \partial_t \eta + \frac{1}{2} \gamma^0 \gamma^i b_i \psi \end{cases}$$

- Carroll fermions can be of **Dirac**, **Majorana**, or **Weyl** spinors

On Supersymmetry

On Supersymmetry (SUSY)

- To supersymmetrize a theory, given by a sum of bosonic and fermionic actions

$$S_{\text{SUSY}} = S_b[\phi] + S_a[A] + S_f[\psi]$$

On Supersymmetry (SUSY)

- To supersymmetrize a theory, given by a sum of bosonic and fermionic actions

$$S_{\text{SUSY}} = S_b[\phi] + S_a[A] + S_f[\psi]$$

one should find “SUSY transformations”:

$$\begin{cases} \delta\phi = \bar{\epsilon}(\dots)\psi \\ \delta\psi = (\dots)\phi\epsilon \end{cases} \quad \text{such that}$$

On Supersymmetry (SUSY)

- To supersymmetrize a theory, given by a sum of bosonic and fermionic actions

$$S_{\text{SUSY}} = S_b[\phi] + S_a[A] + S_f[\psi]$$

one should find “SUSY transformations”:

$$\begin{cases} \delta\phi = \bar{\epsilon}(\dots)\psi \\ \delta\psi = (\dots)\phi\epsilon \end{cases} \quad \text{such that}$$

- leave invariant the “SUSY action”, i.e. $\delta S_{\text{SUSY}} = 0$

On Supersymmetry (SUSY)

- To supersymmetrize a theory, given by a sum of bosonic and fermionic actions

$$S_{\text{SUSY}} = S_b[\phi] + S_a[A] + S_f[\psi]$$

one should find “SUSY transformations”:

$$\begin{cases} \delta\phi = \bar{\epsilon}(\dots)\psi \\ \delta\psi = (\dots)\phi\epsilon \end{cases} \quad \text{such that}$$

1) leave invariant the “SUSY action”, i.e. $\delta S_{\text{SUSY}} = 0$

2) satisfy the “SUSY algebra”, i.e. $[\delta_1, \delta_2] = 2(\bar{\epsilon}_2 \gamma^\mu \epsilon_1) \partial_\mu$

On Supersymmetry (SUSY)

- To supersymmetrize a theory, given by a sum of bosonic and fermionic actions

$$S_{\text{SUSY}} = S_b[\phi] + S_a[A] + S_f[\psi]$$

one should find “SUSY transformations”:

$$\begin{cases} \delta\phi = \bar{\epsilon}(\dots)\psi \\ \delta\psi = (\dots)\phi\epsilon \end{cases} \quad \text{such that}$$

1) leave invariant the “SUSY action”, i.e. $\delta S_{\text{SUSY}} = 0$

2) satisfy the “SUSY algebra”, i.e. $[\delta_1, \delta_2] = 2(\bar{\epsilon}_2 \gamma^\mu \epsilon_1) \partial_\mu$

- In the Carroll supersymmetry, the SUSY algebra reads

Carroll SUSY algebra:
$$[\delta_1, \delta_2] = 2(\bar{\epsilon}_2 \gamma^0 \epsilon_1) \partial_t$$

Carrollian Supersymmetry: $(0, \frac{1}{2})$

Super-electric Carroll theory

We introduce the **super-electric Carroll** action by [K.K, M.N., PRD (2023)]:

$$S_{\text{eC}}^{\text{SUSY}} = \frac{1}{2} \int dt d^3x \left\{ (\partial_t \phi_R)^2 + (\partial_t \phi_I)^2 + F_R^2 + F_I^2 - \bar{\psi} \gamma^0 \partial_t \psi \right\}$$

Super-electric Carroll theory

We introduce the **super-electric Carroll** action by [K.K, M.N., PRD (2023)]:

$$S_{\text{eC}}^{\text{SUSY}} = \frac{1}{2} \int dt d^3x \left\{ (\partial_t \phi_R)^2 + (\partial_t \phi_I)^2 + F_R^2 + F_I^2 - \bar{\psi} \gamma^0 \partial_t \psi \right\}$$

and find that the action is invariant under the Super-electric Carroll transformations

$$\delta \phi_R = \bar{\epsilon} \psi$$

$$\delta \phi_I = \bar{\epsilon} i \gamma^5 \psi$$

$$\delta F_R = -\bar{\epsilon} \gamma^0 \partial_t \psi$$

$$\delta F_I = -\bar{\epsilon} i \gamma^5 \gamma^0 \partial_t \psi$$

$$\delta \psi = \gamma^0 \partial_t (\phi_R + i \gamma^5 \phi_I) \epsilon - (F_R + i \gamma^5 F_I) \epsilon$$

Super-electric Carroll theory

We introduce the **super-electric Carroll** action by [K.K, M.N., PRD (2023)]:

$$S_{\text{eC}}^{\text{SUSY}} = \frac{1}{2} \int dt d^3x \left\{ (\partial_t \phi_R)^2 + (\partial_t \phi_I)^2 + F_R^2 + F_I^2 - \bar{\psi} \gamma^0 \partial_t \psi \right\}$$

and find that the action is invariant under the Super-electric Carroll transformations

$$\begin{aligned} \delta \phi_R &= \bar{\epsilon} \psi \\ \delta \phi_I &= \bar{\epsilon} i \gamma^5 \psi \\ \delta F_R &= -\bar{\epsilon} \gamma^0 \partial_t \psi \\ \delta F_I &= -\bar{\epsilon} i \gamma^5 \gamma^0 \partial_t \psi \\ \delta \psi &= \gamma^0 \partial_t (\phi_R + i \gamma^5 \phi_I) \epsilon - (F_R + i \gamma^5 F_I) \epsilon \end{aligned}$$

We show that these transformations close off-shell for every field

$$[\delta_1, \delta_2] = 2(\bar{\epsilon}_2 \gamma^0 \epsilon_1) \partial_t$$

Super-magnetic Carroll theory

We introduce the **super-magnetic Carroll** action by [K.K, M.N., PRD (2023)]:

$$S_{\text{mC}}^{\text{SUSY}} = \int dt d^3x \left\{ \chi_1 \partial_t \phi_R + \chi_2 \partial_t \phi_I + F_R G_R + F_I G_I + \frac{1}{2} \bar{\lambda} \gamma^0 \psi - \frac{1}{2} \bar{\psi} \gamma^0 \lambda \right\}$$

Super-magnetic Carroll theory

We introduce the **super-magnetic Carroll** action by [K.K, M.N., PRD (2023)]:

$$S_{\text{mC}}^{\text{SUSY}} = \int dt d^3x \left\{ \chi_1 \partial_t \phi_R + \chi_2 \partial_t \phi_I + F_R G_R + F_I G_I + \frac{1}{2} \bar{\lambda} \gamma^0 \psi - \frac{1}{2} \bar{\psi} \gamma^0 \lambda \right\}$$

and find that the action is invariant under the Super-electric Carroll transformations

$$\delta \phi_R = \bar{\epsilon} \psi$$

$$\delta \chi_1 = \bar{\epsilon} \lambda$$

$$\delta F_R = -\bar{\epsilon} \gamma^0 \partial_t \psi$$

$$\delta G_R = -\bar{\epsilon} \gamma^0 \lambda$$

$$\delta \psi = \gamma^0 \partial_t (\phi_R + i \gamma^5 \phi_I) \epsilon - (F_R + i \gamma^5 F_I) \epsilon$$

$$\delta \lambda = \gamma^0 \partial_t (\chi_1 + i \gamma^5 \chi_2) \epsilon - \partial_t (G_R + i \gamma^5 G_I) \epsilon$$

$$\delta \phi_I = \bar{\epsilon} i \gamma^5 \psi$$

$$\delta \chi_2 = \bar{\epsilon} i \gamma^5 \lambda$$

$$\delta F_I = \bar{\epsilon} i \gamma^0 \gamma^5 \partial_t \psi$$

$$\delta G_I = \bar{\epsilon} i \gamma^0 \gamma^5 \lambda$$

Super-magnetic Carroll theory

We introduce the **super-magnetic Carroll** action by [K.K, M.N., PRD (2023)]:

$$S_{\text{mC}}^{\text{SUSY}} = \int dt d^3x \left\{ \chi_1 \partial_t \phi_R + \chi_2 \partial_t \phi_I + F_R G_R + F_I G_I + \frac{1}{2} \bar{\lambda} \gamma^0 \psi - \frac{1}{2} \bar{\psi} \gamma^0 \lambda \right\}$$

and find that the action is invariant under the Super-electric Carroll transformations

$$\delta \phi_R = \bar{\epsilon} \psi$$

$$\delta \chi_1 = \bar{\epsilon} \lambda$$

$$\delta F_R = -\bar{\epsilon} \gamma^0 \partial_t \psi$$

$$\delta G_R = -\bar{\epsilon} \gamma^0 \lambda$$

$$\delta \psi = \gamma^0 \partial_t (\phi_R + i \gamma^5 \phi_I) \epsilon - (F_R + i \gamma^5 F_I) \epsilon$$

$$\delta \lambda = \gamma^0 \partial_t (\chi_1 + i \gamma^5 \chi_2) \epsilon - \partial_t (G_R + i \gamma^5 G_I) \epsilon$$

$$\delta \phi_I = \bar{\epsilon} i \gamma^5 \psi$$

$$\delta \chi_2 = \bar{\epsilon} i \gamma^5 \lambda$$

$$\delta F_I = \bar{\epsilon} i \gamma^0 \gamma^5 \partial_t \psi$$

$$\delta G_I = \bar{\epsilon} i \gamma^0 \gamma^5 \lambda$$

again these transformations close off-shell for every field

$$[\delta_1, \delta_2] = 2(\bar{\epsilon}_2 \gamma^0 \epsilon_1) \partial_t$$

Note

We also derived
Galilei fermions
and generalized them to
Supersymmetry

[K. Koutrolikos, M.Najafzadeh, PRD (2023)]

Conclusion

Summary and open questions

- **In summary:**
 - Carrollian physics can be obtained from a relativistic theory by taking a limit
 - It has usually two sectors, electric and magnetic

Summary and open questions

- **In summary:**

- Carrollian physics can be obtained from a relativistic theory by taking a limit
- It has usually two sectors, electric and magnetic

- **Open questions:**

- vector supermultiplet $(1, \frac{1}{2})$

Summary and open questions

- **In summary:**

- Carrollian physics can be obtained from a relativistic theory by taking a limit
- It has usually two sectors, electric and magnetic

- **Open questions:**

- vector supermultiplet $(1, \frac{1}{2})$
- Carrollian higher spin field theories

Summary and open questions

- **In summary:**

- Carrollian physics can be obtained from a relativistic theory by taking a limit
- It has usually two sectors, electric and magnetic

- **Open questions:**

- vector supermultiplet $(1, \frac{1}{2})$
- Carrollian higher spin field theories
- Conformal Carroll algebras

Summary and open questions

- **In summary:**

- Carrollian physics can be obtained from a relativistic theory by taking a limit
- It has usually two sectors, electric and magnetic

- **Open questions:**

- vector supermultiplet $(1, \frac{1}{2})$
- Carrollian higher spin field theories
- Conformal Carroll algebras
- Carroll/Fracton correspondence

Summary and open questions

- **In summary:**

- Carrollian physics can be obtained from a relativistic theory by taking a limit
- It has usually two sectors, electric and magnetic

- **Open questions:**

- vector supermultiplet $(1, \frac{1}{2})$
- Carrollian higher spin field theories
- Conformal Carroll algebras
- Carroll/Fracton correspondence
- Carroll particle/gravity

Summary and open questions

- **In summary:**

- Carrollian physics can be obtained from a relativistic theory by taking a limit
- It has usually two sectors, electric and magnetic

- **Open questions:**

- vector supermultiplet $(1, \frac{1}{2})$
- Carrollian higher spin field theories
- Conformal Carroll algebras
- Carroll/Fracton correspondence
- Carroll particle/gravity
- ...

Summary and open questions

- **In summary:**

- Carrollian physics can be obtained from a relativistic theory by taking a limit
- It has usually two sectors, electric and magnetic

- **Open questions:**

- vector supermultiplet $(1, \frac{1}{2})$
- Carrollian higher spin field theories
- Conformal Carroll algebras
- Carroll/Fracton correspondence
- Carroll particle/gravity
- ...

Thank you for your attention!